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## Analysis of mathematical modelling of a solar parabolic collector

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### Abstract

The main objective in this paper, a thermal model of a solar parabolic through collector is present. Numerical simulations of the model are performed in Matlab. Temperature evaluation at different nodes is presented and a comparison between a different solar collectors is made.

**Keywords:** Analysis, mathematical, modelling, parabolic, collector

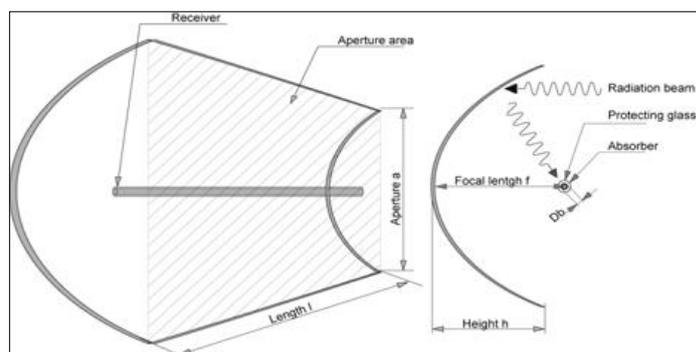
### Introduction

Recently, the environmental consequences have forced many nations in the world to harvest the abundant solar energy for their domestic and industrial applications. Solar thermal systems, in addition to the well-known advantages of renewable resources are very suitable for air-conditioning and refrigeration demands, because solar radiation availability and cooling requirements usually coincide <sup>[1, 2]</sup>.

Solar air-conditioning and refrigeration facilities can also be easily combined with space heating and hot water applications, increasing the yearly solar fraction of buildings. Conventional cold producing machines that are based on vapor compression principle are primary electricity consumers and their working fluids are being banned by international legislation. Solar powered cooling systems as a green cold production technology are the best alternative. Absorption refrigeration is a mature technology that has proved its applicability with the possibility to be driven by low grade solar and waste heat <sup>[3]</sup>.

Xu *et al.* <sup>[4]</sup> presented a study on the comparison of three outdoor test methods for determining the thermal performance of parabolic trough solar collectors. The methods are respectively the steady-state method in the ASHRAE 93 standard, the quasi-dynamic method in the EN 12975-2 standard and a new dynamic method developed by the authors. The comparison shows the advantages and disadvantages of these models according to both the practical operation and weather conditions.

A detailed numerical heat transfer model based on the finite volume method is presented by Hachicha *et al.* <sup>[5]</sup> the different elements of the receiver are discretized into several segments in both axial and azimuthal directions. The set of algebraic equations are solved simultaneously using direct solvers. Results obtained shown a good agreement with experimental data. As shown in Figure 1,



**Fig 1:** Geometry of the solar parabolic trough collector

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$$M_r C_r \frac{dT_r}{dt} = \left( \begin{aligned} &(1-R)G - h_{r-a}(T_r - T_a) \\ &-h_{s-r}(T_r - T_s) - h_{r-g}(T_r - T_g) \end{aligned} \right) \quad (7)$$

Between glass and reflector, it's expressed by

$$M_g C_g \frac{dT_g}{dt} = \alpha_g A_g R G + A_g h_{r-g}(T_r - T_g) - A_{ge}(h_{g-a}(T_g - T_a) + h_{s-g}(T_g - T_s)) + A_{gi}(h_{b-g} + h_d)(T_b - T_g) \quad (8)$$

Similarly by absorber and glass

$$M_b C_b \frac{dT_b}{dt} = \tau_g \alpha_b A_r R G - A_{bi} h_{b-f}(T_b - T_f) - A_{gi}(h_{b-g} + h_d)(T_b - T_g) \quad (9)$$

And finally between HTF and absorber

$$M_f C_f \frac{dT_f}{dt} = A_{pi} h_{p-f}(T_p - T_f) - \dot{m} C_f \frac{dT_f}{dx} \quad (10)$$

With  $M_i$  and  $C_i$  is the mass and specific heat of the  $i$  node material,  $m_f$  is the working fluid mass flow,  $T_s$  the sky temperature calculated by [9]:

$$T_s = 0.0552 T_a^{1.5} \quad (11)$$

Many empirical correlations are used to determine directly the heat transfer coefficients between the different nodes. Reflector and sky [8].

$$h_{s-r} = \sigma \epsilon_r (T_s + T_r)(T_s^2 + T_r^2) \quad (12)$$

Reflector and glass [15, 16].

$$h_{r-g} = \frac{\sigma(T_g^2 + T_r^2)(T_g + T_r)}{\frac{1 - \epsilon_g}{\epsilon_g} + \frac{1}{F_{rg}} + \frac{(1 + \epsilon_r) A_r}{\epsilon_r A_{ge}}} \quad (13)$$

With

$$F_{rg} = \frac{\left[ (W_1 + W_2)^2 + 4 \right]^{\frac{1}{2}} - \left[ (W_2 - W_1)^2 + 4 \right]^{\frac{1}{2}}}{2W_1}$$

$$W_1 = \frac{D_{ge}}{f}$$

$$W_2 = \frac{a}{f}$$

Between reflector and air [10].

$$h_{r-a} = \frac{8.6 I^{0.6}}{l^{0.4}} \quad (14)$$

Between glass and sky [8].

$$h_{s-g} = \sigma \epsilon_g (T_s + T_g)(T_s^2 + T_g^2) \quad (15)$$

Between glass and air [4].

$$h_{g-a} = \frac{4I^{0.58}}{D^{0.42}}$$

Between glass and absorber [6, 8, 11, 12].

$$h_{b-g} = \frac{\sigma(T_g^2 + T_b^2)(T_g + T_b)}{\frac{1 - \epsilon_g}{\epsilon_g} + \frac{1}{F_{gb}} + \frac{(1 + \epsilon_b) A_{gi}}{\epsilon_b A_{be}}} \quad (16)$$

With

$$F_{gb} = \frac{1}{X} \left[ \frac{1}{\pi} \left( \cos^{-1} \left( \frac{B}{A} \right) \frac{1}{2Y} \left( C \cos^{-1} \left( \frac{B}{XA} \right) + B \cos^{-1} \left( \frac{1}{X} \right) - \frac{\pi}{2} A \right) \right) \right]$$

$$X = \frac{D_{gi}}{D_{be}}, \quad Y = \frac{2l}{D_{be}}$$

$$A = X^2 + Y^2 - 1, \quad B = X^2 Y^2 - 1, \quad C = \sqrt{(A+2)^2 - (2X)^2}$$

Between absorber and fluid [13].

For turbulent flow,  $Re > 10^4$ :

$$Nu = 0.125 (0.790 \log(Re) + 1.64)^{-2} Re Pr^{0.34} \quad (18)$$

For laminar flow,  $Re < 10^4$

$$Nu = 3.66 + \frac{0.0668 \left( \frac{D_{bi}}{l} \right) Re Pr}{1 + 0.04 \left[ \left( \frac{D_{bi}}{l} \right) Re Pr \right]^{0.67}} \quad (19)$$

### Solar parabolic trough collector behavior

Programming equations of the parabolic is performed in Matlab. Iterative Gauss method is used to solve these equations after finite difference discretization. The Parabolic trough solar collector unsteady equations are discretized and solved by the finite volume method. A Crank-Nicolson scheme with  $t=1s$  was adopted for the time step.

Calculations are performed using a uniform mesh containing 100 nodes, iterative Gauss method was used and the program was written in Matlab. The collector components' geometrical and thermo physical properties and HTF properties are summarized in Table 1 and Table 2.

Table 1. And 2 illustrate respectively the collector components geometrical and thermal properties and HTF properties.

**Table 1:** Collector properties

	Symbol	Parameter	Value	Unit
<b>Collector</b>	$l$	Collector length	7.800	$m$
	$a$	Aperture	5.000	$m$
	$f$	Focal length	1.840	$m$
<b>Reflector</b>	$Cp_r$	Specific heat	0.581	$kJ/(kg.K)$
	$\rho_r$	Density	2,400	$kg/m^3$
	$e_r$	Thickness	0.005	$m$
	$\varepsilon_r$	Emittance	0.140	–
	$R_r$	Reflectance	0.930	–
<b>Glass</b>	$Cp_g$	Specific heat	1.090	$kJ/(kg.K)$
	$\rho_g$	Density	2,230	$kg/m^3$
	$D_{gi}$	Internal diameter	0.107	$m$
	$D_{ge}$	External diameter	0.114	$m$
	$\tau_g$	Transmittance	0.935	–
	$\varepsilon_g$	Emittance	0.860	–
	$\alpha_g$	Absorptance	0.020	–
<b>Absorber</b>	$Cp_b$	Specific heat	0.500	$kJ/(kg.K)$
	$\rho_b$	Density	8,020	$kg/m^3$
	$D_{bi}$	Internal diameter	0.066	$m$
	$D_{be}$	External diameter	0.0 0	$m$
	$\varepsilon_b$	Emissivity	0.140	–
	$\alpha_b$	Absorptance	0.905	–
	$p_v$	Vacuum pressure	0.010	$bar$

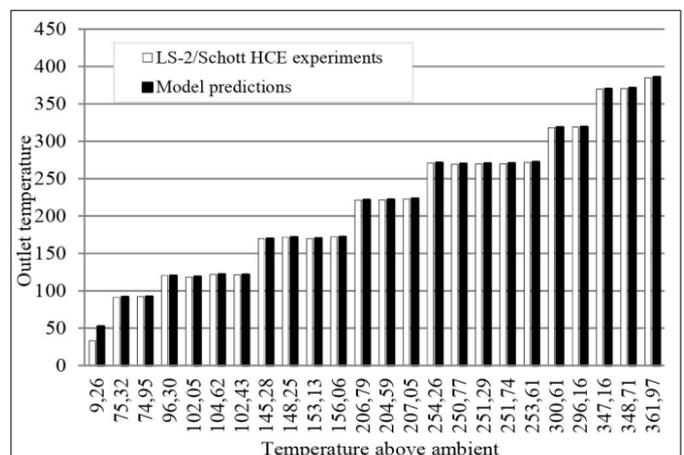
**Table 2:** Fluid Properties

Parameter	Value	Unit
Specific heat	1,616	$kJ/(kg.K)$
Density	935	$kg/m^3$
Conductivity	0.134	$W/(m.K)$
Viscosity	9.1	$Pa.s$
Critical temperature	367	$^{\circ}C$
Critical pressure	10.7	$bar$
Critical density	3.22	$l/kg$

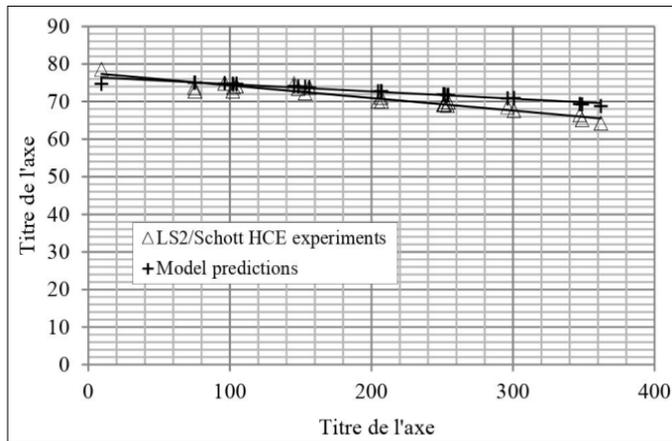
**Result and Discussion**

According to Figure 4 shows a comparison of the evolution the outlet HTF temperature with respect to the temperature above ambient between model predictions and test data of the LS-2/Schott HCE. Except the first value, a good

agreement is found that the difference between predicted and experimental values of temperature does not exceed one degree (1K).



**Fig 4:** Outlet HTF temperature vs. temperature above ambient

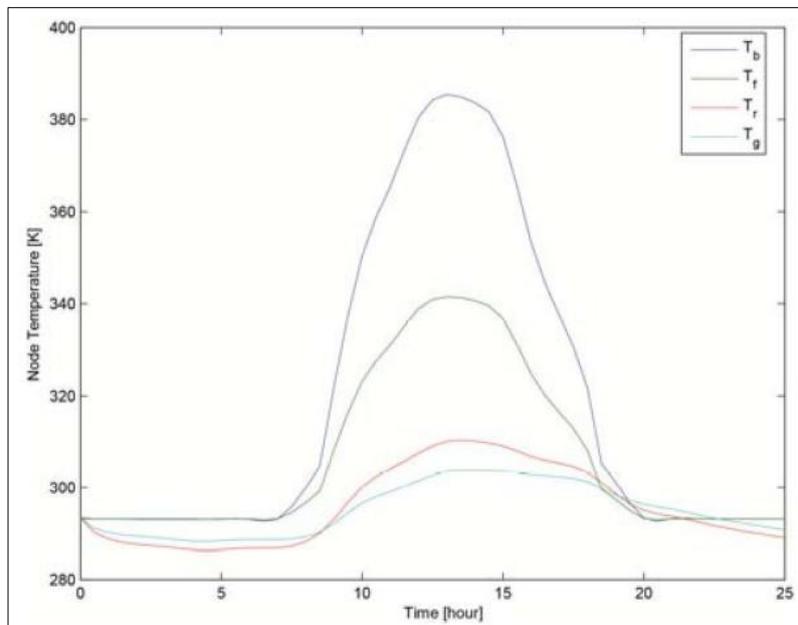


**Fig 5:** Peak efficiencies vs. average HTF temperature through the HCE above ambient

Similarly, figure 5 shows a comparison between predicted and measured peak efficiencies versus the HTF temperature above ambient. Here again, a good agreement is found. Nevertheless, a maximum prediction error of 6.56% is observed between predicted and measured values changes according to the temperature to obtain

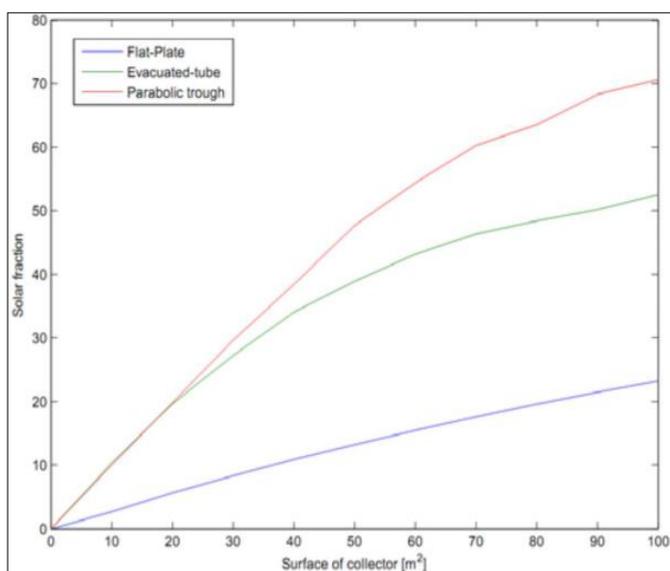
These results allow for the use of the model for predicting the approximate temperature of the fluid outlet and the energy required to activate desorption of the binary solution in the absorption machine.

Simulations allow studying the temperature evolution and hence the effectiveness of the collector at any time of the day throughout the year. Temperature evolution at each node is in a good agreement with the studied literature. Absorber temperature is greater compared to other temperatures due to its position in the focal line and its absorptivity, as shown in Figure 6.



**Fig 6:** Daily Temperature evolution at the different nodes

Simulations allow also studying the efficiency of some collectors and to detect the most suitable energy solution. As shown in Figure 7



**Fig 7:** Solar fraction for different collector types

Illustrates a comparison of collectors' solar fractions relative to the surface and allows choosing the most appropriate collector for the solar cooling installation, thus optimizing its surface. It shows that the flat plate collector has the lowest solar fraction compared to the two other collectors. For the two others, the preference depends on climatic conditions and surface. For small areas, the evacuated tube collector is more convenient. But when using larger areas, the concentration collector becomes the most reliable.

**Conclusion**

The present paper phasing between the need for cold and Solar gain in generally excellent, with sometimes a slight delay which can be compensated by storage. The use of solar energy for air-conditioning is a promising alternative.

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